Shaping Jets with the Ambient Medium

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Jets = highly energetic outflows



Observations show jet regime changes



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Cutting-edge numerical simulation

GPU-accelerated, 3D General Relativistic Magnetohydrodynamic code, *HAMR*

Magnetized torus launches powerful jets



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Disk winds create funnel near black hole





Pressure balance between disk and jet creates parabolic profile

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Probing effects of the ambient medium

Infinite Ambient Medium

Cutoff Ambient Medium



Infinite Ambient Medium

Cutoff Ambient Medium



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Rapid expansion of magnetization

R A

Infinite Ambient Medium



Cutoff Ambient Medium



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Jet shape in simulations



Infinite medium creates cylindrical jets



Jet regime changes drastically near cutoff



Expanding jets speed up



Takeaways

Disk winds shape jet into a parabola near the black hole.

A r^{-1} ambient density profile can provide sufficient pressure to **collimate jets into cylinders**.

Escaping regions of higher density allows for transition to a **conical free expansion regime**.

Future work:

- ✤ larger scale separations
- different ambient medium profiles

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Resolving jets at large scale separation



Base grid is (640 x 288 x 288) in spherical coordinates

Adaptive Mesh Refinement allows jet to be resolved at large distances

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How to pick out jets



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R A

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Jet power



Jet efficiency



Magnetic flux on the black hole



Mass accretion rate

