#### Multiscale Problems in Fluctuating Hydrodynamics

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Giant Fluctuations

#### Fractal Fronts in Diffusive Mixing



Snapshots of concentration in a miscible mixture showing the development of a *rough* diffusive interface between two miscible fluids in zero gravity [1, 2, 3]. A similar pattern is seen over a broad range of Schmidt numbers and is affected strongly by nonzero gravity.

#### **Giant Fluctuations**

#### Fluctuating Navier-Stokes Equations

- We will consider a binary fluid mixture with mass concentration  $c = \rho_1/\rho$  for two fluids that are dynamically identical, where  $\rho = \rho_1 + \rho_2$  (e.g., fluorescently-labeled molecules).
- Ignoring density and temperature fluctuations, equations of incompressible isothermal fluctuating hydrodynamics are

$$\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v} = -\nabla \pi + \nu \nabla^2 \mathbf{v} + \nabla \cdot \left(\sqrt{2\nu\rho^{-1} k_B T} \mathcal{W}\right)$$
$$\partial_t c + \mathbf{v} \cdot \nabla c = \chi \nabla^2 c + \nabla \cdot \left(\sqrt{2m\chi\rho^{-1} c(1-c)} \mathcal{W}^{(c)}\right),$$

where the **kinematic viscosity**  $\nu = \eta/\rho$ , and  $\pi$  is determined from incompressibility,  $\nabla \cdot \mathbf{v} = 0$ .

 We assume that *W* can be modeled as spatio-temporal white noise (a delta-correlated Gaussian random field), e.g.,

$$\langle \mathcal{W}_{ij}(\mathbf{r},t)\mathcal{W}_{kl}^{\star}(\mathbf{r}',t')\rangle = (\delta_{ik}\delta_{jl}+\delta_{il}\delta_{jk})\,\delta(t-t')\delta(\mathbf{r}-\mathbf{r}').$$

#### Nonequilibrium Fluctuations

- When macroscopic gradients are present, steady-state thermal fluctuations become **long-range correlated**.
- Consider a binary mixture of fluids and consider concentration fluctuations around a steady state c<sub>0</sub>(r):

$$c(\mathbf{r},t) = c_0(\mathbf{r}) + \delta c(\mathbf{r},t)$$

• The concentration fluctuations are advected by the random velocities  $\mathbf{v}(\mathbf{r}, t) = \delta \mathbf{v}(\mathbf{r}, t)$ , approximately:

$$\partial_t \left( \delta c \right) + \left( \delta \mathbf{v} \right) \cdot \boldsymbol{\nabla} c_0 = \chi \boldsymbol{\nabla}^2 \left( \delta c \right) + \sqrt{2 \chi k_B T} \left( \boldsymbol{\nabla} \cdot \boldsymbol{\mathcal{W}}_c \right)$$

• The velocity fluctuations drive and amplify the concentration fluctuations leading to so-called **giant fluctuations** [2].

#### Back of the Envelope

• The coupled *linearized velocity*-concentration system in **one dimension**:

$$\begin{aligned} \mathbf{v}_t &= \nu \mathbf{v}_{\mathsf{X}\mathsf{X}} + \sqrt{2\nu} \, W_{\mathsf{X}} \\ \mathbf{c}_t &= \chi \mathbf{c}_{\mathsf{X}\mathsf{X}} - \nu \bar{\mathbf{c}}_{\mathsf{X}}, \end{aligned}$$

where  $g = \bar{c}_x$  is the imposed background concentration gradient.

• The linearized system can be easily solved in Fourier space to give a **power-law divergence** for the spectrum of the concentration fluctuations as a function of wavenumber *k*,

$$\langle \hat{c}\hat{c}^{\star}
angle \sim rac{\left(ar{c}_{x}
ight)^{2}}{\chi(\chi+
u)k^{4}}.$$

- Concentration fluctuations become **long-ranged** and are enhanced as the square of the gradient, to values much larger than equilibrium fluctuations.
- In real life the divergence is **suppressed** by surface tension, gravity, or boundaries (usually in that order).

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#### Giant Fluctuations

### Diffusive Mixing in Gravity



#### **Giant Fluctuations**

#### Giant Fluctuations in Experiments



Experimental results by A. Vailati *et al.* from a microgravity environment [2] showing the enhancement of concentration fluctuations in space (box scale is **macroscopic**: 5mm on the side, 1mm thick).

#### Low Mach Approximation

For isothermal mixtures of fluids with unequal densities, the incompressible approximation needs to be replaced with a **low Mach approximation** 

$$D_{t}\rho = -\rho \left( \boldsymbol{\nabla} \cdot \boldsymbol{v} \right)$$
  
$$\rho \left( D_{t} \boldsymbol{v} \right) = -\boldsymbol{\nabla} P + \boldsymbol{\nabla} \cdot \left[ \eta \left( \boldsymbol{\nabla} \boldsymbol{v} + \boldsymbol{\nabla} \boldsymbol{v}^{T} \right) + \boldsymbol{\Sigma} \right]$$
  
$$\rho \left( D_{t} c \right) = \boldsymbol{\nabla} \cdot \left[ \rho \chi \left( \boldsymbol{\nabla} c \right) + \boldsymbol{\Psi} \right],$$

where  $D_t \Box = \partial_t \Box + \mathbf{v} \cdot \nabla(\Box)$  and  $\boldsymbol{\Sigma}$  and  $\boldsymbol{\Psi}$  are stochastic fluxes determined from fluctuation-dissipation balance.

The incompressibility condition is replaced by the equation of state (EOS) constraint

$$\boldsymbol{\nabla} \cdot \mathbf{v} = \rho^{-1} \left( \frac{\partial \rho}{\partial c} \right)_{P,T} (D_t c).$$

# Fluctuating Hydrodynamics Equations

- Adding stochastic fluxes to the **non-linear** NS equations produces **ill-behaved stochastic PDEs** (solution is too irregular).
- No problem if we **linearize** the equations around a **steady mean state**, to obtain equations for the fluctuations around the mean.
- Finite-volume discretizations naturally impose a grid-scale **regularization** (smoothing) of the stochastic forcing.
- A renormalization of the transport coefficients is also necessary [1].
- We have algorithms and codes to solve the compressible equations (collocated and staggered grid), and recently also the incompressible and low Mach number ones (staggered grid) [4, 3].
- Solving these sort of equations numerically requires paying attention to **discrete fluctuation-dissipation balance**, in addition to the usual deterministic difficulties [4].

### Finite-Volume Schemes

$$c_t = -\mathbf{v} \cdot \nabla c + \chi \nabla^2 c + \nabla \cdot \left(\sqrt{2\chi} \mathcal{W}\right) = \nabla \cdot \left[-c\mathbf{v} + \chi \nabla c + \sqrt{2\chi} \mathcal{W}\right]$$

• Generic finite-volume spatial discretization

$$\mathbf{c}_{t} = \mathbf{D}\left[ \left( -\mathbf{V}\mathbf{c} + \mathbf{G}\mathbf{c} \right) + \sqrt{2\chi/\left(\Delta t \Delta V\right)} \mathbf{W} \right],$$

where D : faces  $\rightarrow$  cells is a **conservative** discrete divergence, G : cells  $\rightarrow$  faces is a discrete gradient.

- Here **W** is a collection of random normal numbers representing the (face-centered) stochastic fluxes.
- The divergence and gradient should be duals,  $D^* = -G$ .
- Advection should be **skew-adjoint** (non-dissipative) if  $\nabla \cdot \mathbf{v} = 0$ ,

$$(DV)^* = -(DV)$$
 if  $(DV) \mathbf{1} = \mathbf{0}$ .

- We performed event-driven **hard disk simulations** of diffusive mixing with about 1.25 million disks.
- The two species had equal molecular diameter but potentially different molecular masses, with density ratio  $R = m_2/m_1 = 1, 2$  or 4.
- In order to convert the particle data to hydrodynamic data, we employed finite-volume averaging over a grid of  $128^2$  hydrodynamic cells  $10 \times 10$  molecular diameters (about 76 disks per hydrodynamic cell).
- We also performed fluctuating low Mach number **finite-volume simulations** using the same grid of hydrodynamic cells, at only a small fraction of the computational cost [5].
- Quantitative statistical comparison between the molecular dynamics and fluctuating hydrodynamics was excellent once the values of the **bare diffusion** and **viscosity** were adjusted based on the level of coarse-graining.

Fluctuating Hydrodynamics

### Hard-Disk Simulations



Fluctuating Hydrodynamics

### "Hard-Sphere" Simulations



Limiting Diffusive Dynamics

### Passively-Advected (Fluorescent) Tracers



#### Diffusion by Velocity Fluctuations

• Consider a large collection of **passively-advected particles** immersed in a fluctuating Stokes velocity field,

$$\partial_t \mathbf{v} = \mathcal{P} \left[ \nu \nabla^2 \mathbf{v} + \nabla \cdot \left( \sqrt{2\nu\rho^{-1} k_B T} \mathcal{W} \right) \right]$$
$$\partial_t c = -\mathbf{v} \cdot \nabla c + \chi \nabla^2 c + \nabla \cdot \left( \sqrt{2\chi c} \mathcal{W}^{(c)} \right),$$

where c is the number density for the particles, and  $\mathcal{P}$  is the orthogonal projection onto the space of divergence-free velocity fields.

• In liquids diffusion of mass is much slower than diffusion of momentum,  $\chi \ll \nu$ , leading to a **Schmidt number** 

$$S_c = rac{
u}{\chi} \sim 10^3.$$

• [With *Eric Vanden-Eijnden*]: There exists a limiting dynamics for c in the limit  $S_c \rightarrow \infty$  in the scaling

$$u = \chi S_c, \quad \chi(\chi + \nu) \approx \chi \nu = \text{const}$$

#### Rescaling Dynamics

• Consider a family of equations with rescaled coefficients

$$\nu' = \epsilon^{-1}\nu, \quad \chi' = \epsilon\chi,$$

which has  $\nu'\chi' = \chi\nu$  but  $S_c' = \epsilon^{-2}S_c$ .

- For  $\epsilon = 1$  we get the original dynamics, and as  $\epsilon \to 0$  we get the limiting dynamics  $S_c \to \infty$ .
- Rescale time as  $t' = e^{-1}t$ , to get the rescaled equations

$$\begin{aligned} \partial_{t'} \mathbf{v} = \mathcal{P} \left[ \epsilon^{-2} \nu \nabla^2 \mathbf{v} + \nabla \cdot \left( \sqrt{2\epsilon^{-2} \nu \rho^{-1} k_B T} \mathcal{W} \right) \right] \\ \partial_{t'} c = -\epsilon^{-1} \mathbf{v} \cdot \nabla c + \chi \nabla^2 c + \text{stoch.} \end{aligned}$$

 On the rescaled time scale the dynamics will be very similar to the limiting dynamics if 
 *e* is small, specifically, if there is a very large separation of scales between the velocity and concentration dynamics.

#### Adiabatic Elimination of $\mathbf{v}$

- The existence of the limiting dynamics follows from well-established limit theorems, see review by Eric Vanden-Eijnden in Section IV of [6] or a series of three articles on "Adiabatic elimination in stochastic systems" in Phys. Rev. A [7].
- Briefly, if the Liouville operator has the form

$$L = L_0 + \epsilon^{-1}L_1 + \epsilon^{-2}L_2,$$

in the limiting dynamics  $\epsilon \rightarrow {\rm 0}$  we have

$$L' = \mathcal{P}_{\epsilon} L_0 \mathcal{P}_{\epsilon} - \mathcal{P}_{\epsilon} L_1 L_2^{-1} L_1 \mathcal{P}_{\epsilon},$$

where  $\mathcal{P}_{\epsilon}$  is a Zwanzig projection operator  $\bullet$  In our specific case

$$L_0 \longleftrightarrow \chi \nabla^2 c, \quad L_1 \longleftrightarrow -\epsilon^{-1} \mathbf{v} \cdot \nabla c, \quad L_2 \longleftrightarrow \mathcal{P} \epsilon^{-2} \nu \nabla^2 \mathbf{v}$$

## Limiting Dynamics

• A Fourier-space calculation gives *approximately* the following limiting **stochastic advection-diffusion equation** for concentration (common in turbulence models):

$$\partial_t c = -\mathbf{v} \cdot \nabla c + (\chi + \Delta \chi) \nabla^2 c,$$

where  $\Delta \chi$  is a **renormalization** of the diffusion coefficient [1], approximated here by a local diffusion.

• The advection velocity here is a **white-in-time** process that can be sampled by solving the steady Stokes equation

$$\nabla \pi = \nu \nabla^2 \mathbf{v} + \nabla \cdot \left( \sqrt{2\nu \rho^{-1} \, k_B T} \, \mathcal{W} \right)$$
$$\nabla \cdot \mathbf{v} = 0.$$

#### Limiting Diffusive Dynamics

### Simulating the Limiting Dynamics

The limiting dynamics can be efficiently simulated using the following **predictor-corrector algorithm**:

Generate a random advection velocity

$$\nabla \pi^{n+\frac{1}{2}} = \nu \left( \nabla^2 \mathbf{v}^n \right) + \Delta t^{-\frac{1}{2}} \nabla \cdot \left( \sqrt{2\nu \rho^{-1} k_B T} \, \mathcal{W}^n \right)$$
$$\nabla \cdot \mathbf{v}^n = 0.$$

Itake a predictor step for concentration, e.g., using Crank-Nicolson,

$$\frac{\tilde{c}^{n+1}-c^n}{\Delta t}=-\mathbf{v}^n\cdot\boldsymbol{\nabla}c^n+\chi\boldsymbol{\nabla}^2\left(\frac{c^n+\tilde{c}^{n+1}}{2}\right).$$

Take a corrector step for concentration

$$\frac{c^{n+1}-c^n}{\Delta t} = -\mathbf{v}^n \cdot \nabla\left(\frac{c^n + \tilde{c}^{n+1}}{2}\right) + \chi \nabla^2\left(\frac{c^n + c^{n+1}}{2}\right)$$

#### Multiscale Temporal Integrators

• There are two ways in which the rescaled dynamics can be used. Recall that

$$\nu' = \epsilon^{-1} \nu, \quad \chi' = \epsilon \chi \quad \Rightarrow \quad S'_c = \epsilon^{-2} S_c.$$

- The first one is to increase e ≫ 1 and use this to decrease the separation of time scales to the point where the rescaled dynamics is computationally feasible, in the spirit of the seamless HMM method (Vanden-Eijnden et al).
- The second one is to **decrease**  $\epsilon \to 0$  and directly simulate the liming dynamics, which assumes infinite separation of scales (overdamped limit).
- If there is strong separation of scales in the original problem either will do and in fact there may exist integrators that can handle the limit *ϵ* → 0 gracefully (stiffly-accurate integrators).

Limiting Diffusive Dynamics

#### Changing $S_c$ from 1 to $\infty$



#### Questionable Separation

- The above animation makes it clear  $S_c$  needs to be very large to be close to the limiting dynamics.
- The separation of time scales between the slowest velocity mode and the fastest concentration mode is

$$\frac{k_{\max}^2\nu}{k_{\min}^2\chi} = \frac{S_c}{N_c^2},$$

where  $N_c$  is the number of modes (along a direction).

- Full separation of scales requires  $S_c \gg N_c^2$ , which is often not met in practice, e.g.,  $S_c \sim 500$  in a typical liquid like water.
- Similarly **questionable** is the **assumption** that particles immersed in a fluid follow a diffusion equation: what about large-scale slow velocity fluctuations?
- Under certain conditions the limiting dynamics should be a good approximation, but seems hard to justify in general.

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#### Conclusions

- Fluctuations are **not just a microscopic phenomenon**: giant fluctuations can reach macroscopic dimensions or certainly dimensions much larger than molecular.
- Fluctuating hydrodynamics agrees with molecular dynamics of diffusive mixing in mixtures of hard disks and seems to be a very good coarse-grained model for fluids, despite unresolved issues.
- Diffusion is strongly affected and often dominated by **advection by velocity fluctuations**.
- Even coarse-grained methods need to be accelerated due to **large separation of time scales** between advective and diffusive phenomena. One can both decrease or increase the separation of scales to allow for efficient simulation.
- In the case of SPDEs there are many (≫ 1) length and time scales.
   Can one construct many-scale temporal integrators that are accurate even when they under-resolve the fast fluctuations?

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